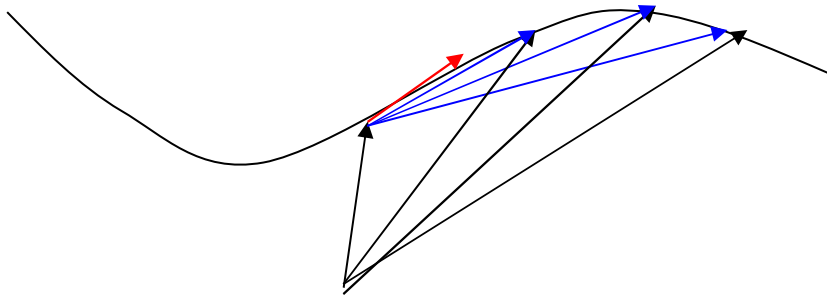


Proprieta' intrinseche di velocita' e accelerazione
Posizione del punto: se la traiettoria e' nota, si puo' dare anche con ascissa curvilinea lungo la traiettoria

$$s = s(t)$$

$$v = \frac{ds}{dt} \text{ vel. scalare}$$

Esaminando la figura:



$$d\mathbf{r} = ds\hat{\mathbf{u}}_T$$

$\hat{\mathbf{u}}_T$ versore tangente alla traiettoria ($\leftarrow non$ costante)

$$\rightarrow \mathbf{v} = \frac{d\mathbf{r}}{dt} = \frac{ds}{dt}\hat{\mathbf{u}}_T = v\hat{\mathbf{u}}_T$$

Scritta in forma indipendente dalle componenti

Piu' formalmente:

$$\mathbf{v} = \lim_{\Delta t \rightarrow 0} \frac{\Delta \mathbf{r}}{\Delta t} = \lim_{\substack{\Delta t \rightarrow 0 \\ \Delta s \rightarrow 0}} \frac{\Delta \mathbf{r}}{\Delta s} \frac{\Delta s}{\Delta t}$$

$$\lim_{\Delta s \rightarrow 0} \frac{\Delta \mathbf{r}}{\Delta s} = \hat{\mathbf{u}}_T$$

$$\rightarrow \mathbf{v} = \frac{ds}{dt} \hat{\mathbf{u}}_T$$

Accelerazione

$$\mathbf{v} = \frac{ds}{dt} \hat{\mathbf{u}}_T = v \hat{\mathbf{u}}_T$$

$$\rightarrow \mathbf{a} = \frac{d\mathbf{v}}{dt} = \frac{d}{dt} (v \hat{\mathbf{u}}_T) = \frac{dv}{dt} \hat{\mathbf{u}}_T + v \frac{d\hat{\mathbf{u}}_T}{dt}$$

Usando componenti cartesiane di $\hat{\mathbf{u}}_T$:

φ angolo rispetto ad asse x

$$\hat{\mathbf{u}}_T = \cos \varphi \hat{\mathbf{u}}_x + \sin \varphi \hat{\mathbf{u}}_y$$

$$\rightarrow \frac{d\hat{\mathbf{u}}_T}{dt} = -\sin \varphi \frac{d\varphi}{dt} \hat{\mathbf{u}}_x + \cos \varphi \frac{d\varphi}{dt} \hat{\mathbf{u}}_y$$

$$\hat{\mathbf{u}}_N = -\sin \varphi \hat{\mathbf{u}}_x + \cos \varphi \hat{\mathbf{u}}_y$$

$$\hat{\mathbf{u}}_N \cdot \hat{\mathbf{u}}_T = 0 \rightarrow \hat{\mathbf{u}}_N \perp \hat{\mathbf{u}}_T \text{ vers. normale}$$

$\hat{\mathbf{u}}_N$ diretto verso il centro istantaneo di curvatura

$$\rightarrow \frac{d\hat{\mathbf{u}}_T}{dt} = \frac{d\varphi}{dt} \hat{\mathbf{u}}_N$$

$$\rightarrow \mathbf{a} = \frac{dv}{dt} \hat{\mathbf{u}}_T + v \frac{d\varphi}{dt} \hat{\mathbf{u}}_N$$

Introduciamo:

$R =$ raggio di curvatura punto per punto

$$\rightarrow ds = R d\varphi$$

$$\rightarrow \frac{d\varphi}{ds} = \frac{1}{R}$$

$$\frac{d\varphi}{dt} = \frac{d\varphi}{ds} \frac{ds}{dt} = \frac{1}{R} v$$

$$\rightarrow \mathbf{a} = \frac{dv}{dt} \hat{\mathbf{u}}_T + v \frac{d\varphi}{dt} \hat{\mathbf{u}}_N = \frac{dv}{dt} \hat{\mathbf{u}}_T + \frac{v^2}{R} \hat{\mathbf{u}}_N$$

Acc. tangenziale e normale/centripeta

$$\rightarrow a = \sqrt{\left(\frac{dv}{dt}\right)^2 + \frac{v^4}{R^2}}$$

Riassumendo: Derivate dei versori T e N rispetto a un parametro della traiettoria

$$\frac{d\hat{\mathbf{u}}_T}{dt} = \frac{d\hat{\mathbf{u}}_T}{ds} \frac{ds}{dt} = \frac{d\varphi}{ds} \frac{ds}{dt} \hat{\mathbf{u}}_N$$

$$\rightarrow \begin{cases} \frac{d\hat{\mathbf{u}}_T}{ds} = \frac{d\varphi}{ds} \hat{\mathbf{u}}_N = \frac{1}{R} \hat{\mathbf{u}}_N \\ \frac{d\hat{\mathbf{u}}_N}{ds} = -\frac{1}{R} \hat{\mathbf{u}}_T \end{cases}$$

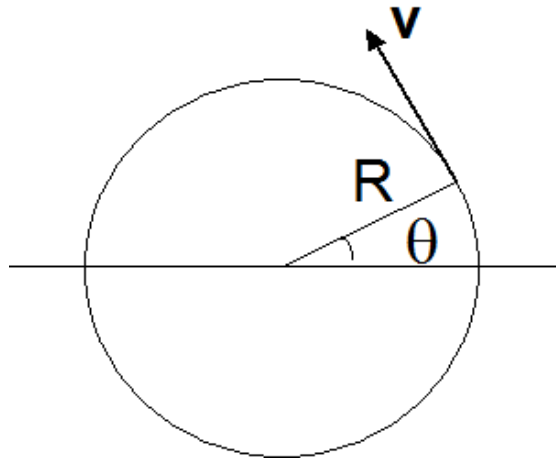
Formule di Frenet-Serret

$\frac{1}{R}$ curvatura ($rad\ m^{-1}$) della traiettoria

Moto circolare

Traiettoria: circonferenza

Nel piano del moto:



$$\mathbf{v} = \frac{d\mathbf{r}}{dt} = \frac{ds}{dt} \hat{\mathbf{u}}_T$$

$$s = R\theta$$

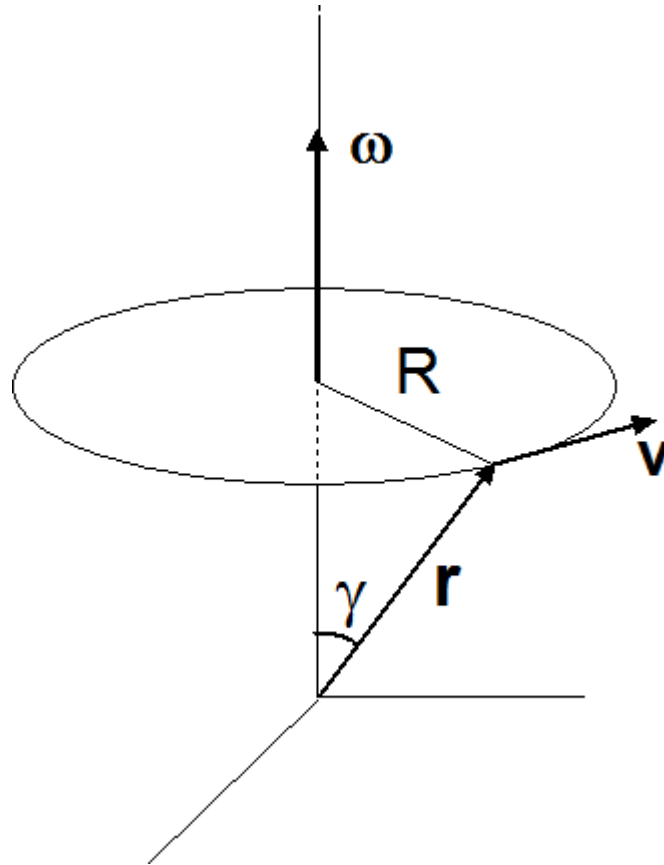
$$\rightarrow \mathbf{v} = \frac{ds}{dt} \hat{\mathbf{u}}_T = R \frac{d\theta}{dt} \hat{\mathbf{u}}_T$$

$$\omega = \frac{d\theta}{dt} \text{ vel. angolare}$$

$$\rightarrow v = \omega R$$

$$\mathbf{v} \perp \mathbf{r}$$

Misurando le posizioni rispetto a un'origine qualsiasi sull'asse di rotazione:



$$R = r \sin \gamma$$

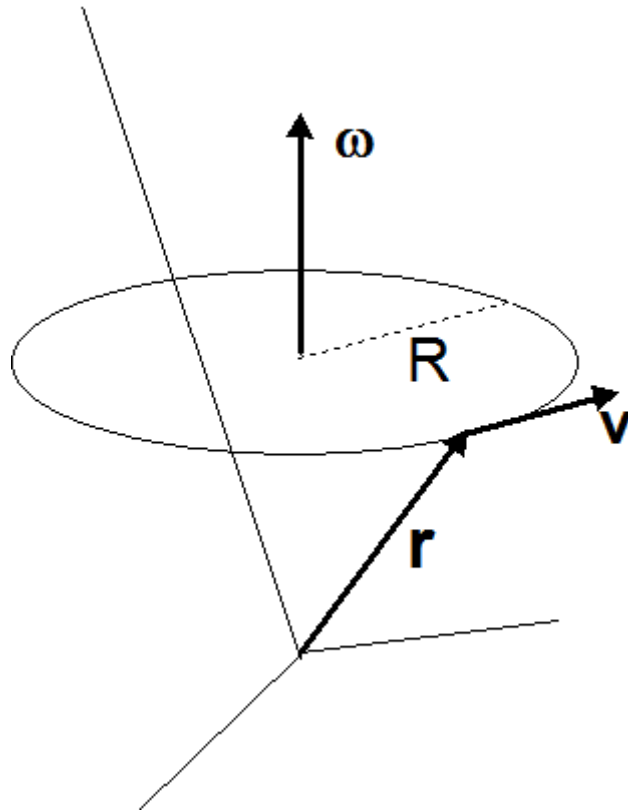
$$\rightarrow v = \omega R \sin \gamma$$

$$\rightarrow |v| = |\omega \times r|$$

$$v \perp r, v \perp \omega$$

$$\rightarrow v = \omega \times r$$

Relazione fra vettori → Indipendente da scelta assi
→Valida in generale



Acc. Angolare:

$$\alpha = \frac{d\omega}{dt}$$

Se:

$$\alpha = \frac{d\omega}{dt} \parallel \omega :$$

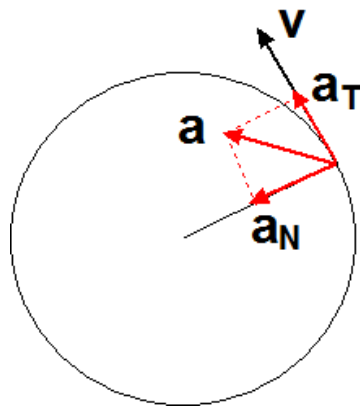
→ Eq. oraria per θ , velocità'

$$\left. \begin{aligned} \theta(t) &= \theta_0 + \int_{t_0}^t \omega(t') dt' \\ \omega(t) &= \omega_0 + \int_{t_0}^t \alpha(t') dt' \end{aligned} \right\} \text{ analogo a moto rettilineo}$$

Acc. tangenziale e normale:

$$a_T = \frac{dv}{dt} = R \frac{d\omega}{dt} = R\alpha$$

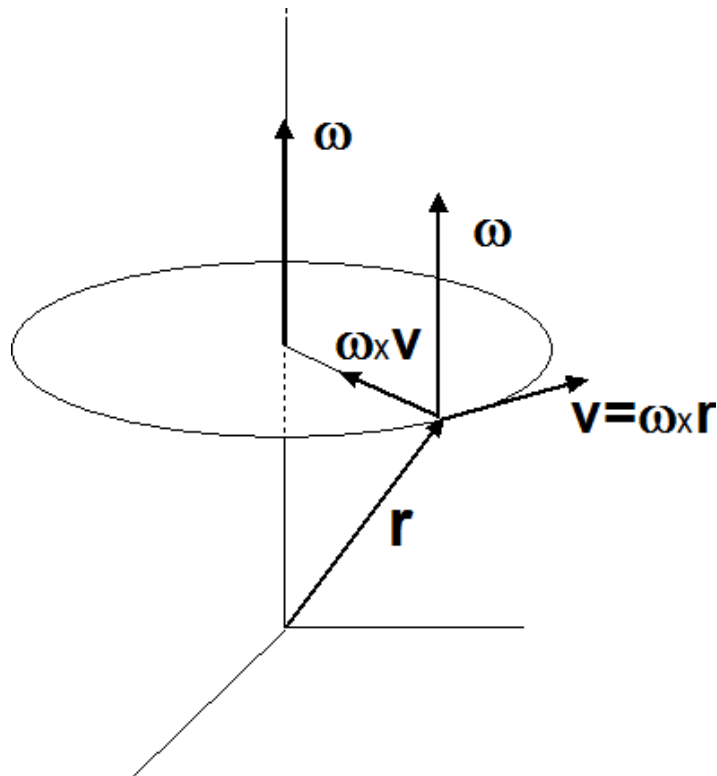
$$a_N = \frac{v^2}{R} = \omega^2 R$$



Moto circolare uniforme

$$\mathbf{a} = \frac{d\mathbf{v}}{dt} = \frac{d}{dt}(\boldsymbol{\omega} \times \mathbf{r})$$

$$\boldsymbol{\omega} = \text{costante} \rightarrow \mathbf{a} = \frac{d}{dt}(\boldsymbol{\omega} \times \mathbf{r}) = \boldsymbol{\omega} \times \underbrace{\frac{d\mathbf{r}}{dt}}_{\mathbf{v}} = \boldsymbol{\omega} \times (\boldsymbol{\omega} \times \mathbf{r})$$



ω frequenza angolare = pulsazione rads^{-1}

$\nu = \frac{\omega}{2\pi}$ frequenza Hz

$T = \frac{1}{\nu} = \frac{2\pi}{\omega}$ periodo s

Velocita' e accelerazione su una ruota di bicicletta

$$v = 25 \text{ km h}^{-1} \sim 7 \text{ m s}^{-1}$$

$$R = 40 \text{ cm} = 0.4 \text{ m}$$

$$\rightarrow \omega \sim \frac{v}{R} = \frac{7}{0.4} \text{ rad s}^{-1} = 17.5 \text{ rad s}^{-1}$$

$$\rightarrow a = \omega^2 R = 122.5 \text{ m s}^{-2}$$

c.s. sulla curva di un'autostrada

$$v = 100 \text{ km h}^{-1} \sim 28 \text{ m s}^{-1}$$

$$R = 1 \text{ km} = 10^3 \text{ m}$$

$$\rightarrow \omega \sim \frac{v}{R} = \frac{28}{10^3} \text{ rad s}^{-1} = 2.8 \cdot 10^{-2} \text{ rad s}^{-1}$$

$$\rightarrow a = \omega^2 R = 7.8 \cdot 10^{-4} \cdot 10^3 = 0.78 \text{ m s}^{-2}$$

c.s. alla superficie della Terra (sull'Equatore)

$$v = 40000 / 24 \text{ km h}^{-1} \sim 463 \text{ m s}^{-1}$$

$$R = 6400 \text{ km} = 6.4 \cdot 10^6 \text{ m}$$

$$\rightarrow \omega \sim \frac{v}{R} = \frac{463}{6.4 \cdot 10^6} \text{ rad s}^{-1} = 7.2 \cdot 10^{-5} \text{ rad s}^{-1}$$

$$\rightarrow a = \omega^2 R = 3.32 \cdot 10^{-2} \text{ m s}^{-2}$$

c.s. per il moto orbitale della Terra

$$v = 9.42 \cdot 10^{11} / 3.15^7 \text{ ms}^{-1} \sim 2.99 \cdot 10^4 \text{ m s}^{-1}$$

$$R = 1.5 \cdot 10^8 \text{ km} = 1.5 \cdot 10^{11} \text{ m}$$

$$\rightarrow \omega \sim \frac{v}{R} = \frac{2.99 \cdot 10^4}{1.5 \cdot 10^{11}} \text{ rad s}^{-1} = 1.98 \cdot 10^{-7} \text{ rad s}^{-1}$$

$$\rightarrow a = \omega^2 R = 4 \cdot 10^{-14} \cdot 1.5 \cdot 10^{11} = 6 \cdot 10^{-3} \text{ m s}^{-2}$$

c.s. per il moto orbitale della Galassia

$$d_{sole} \approx 3.268000 AL = 2.610^4 AL$$

$$1AL = 310^8 ms^{-1} 7.5710^8 s = 2.2710^{17} m$$

$$2\pi d_{sole} = 2\pi 5.910^{21} m = 3.7110^{22} m$$

$$T = 220 \cdot 10^6 \text{anni} = 220 \cdot 7.5710^{14} s = 1.6610^{17} s$$

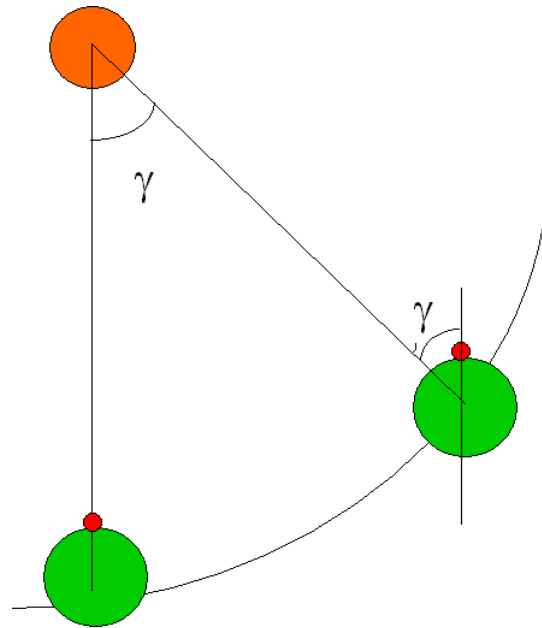
$$v = 3.7110^{22} m / 1.6610^{17} s \sim 2.2410^5 m s^{-1}$$

$$R = 5.910^{21} m$$

$$\rightarrow \omega \sim \frac{v}{R} = \frac{2.2410^5}{5.910^{21}} rad s^{-1} = 0.3810^{-16} rad s^{-1}$$

$$\rightarrow a = \omega^2 R = 0.1410^{-32} 5.910^{21} = 0.8310^{-11} m s^{-2}$$

Es: Vel. angolare della Terra, giorno solare, giorno sidereo



$$\omega_{rot} = \frac{2\pi}{T_{sid}} \approx \frac{2\pi}{T_{sol}} \approx \frac{2\pi}{86400} = 7.27210^{-5} rad s^{-1}$$

$$\Delta t = \frac{ang(1 \text{ day})}{\omega_{rot}} \approx \frac{1.710^{-2}}{7.27210^{-5}} \approx 234 s$$